

# インジウムを用いた太陽 ニュートリノ半導体検出器 の開発XI

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宮城教育大学 福田善之、加賀谷亮太

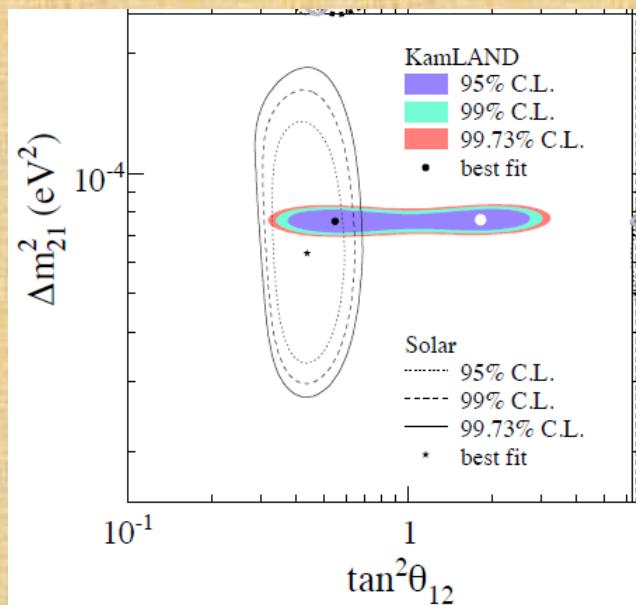
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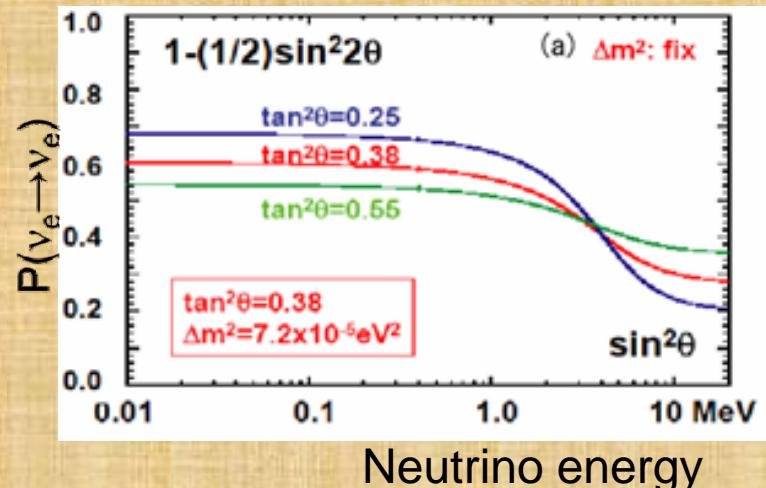
# Motivation

Allowed region obtained by combined results and KamLAND



Survival probability for solar matter oscillation below 1MeV

[Y.Suzuki@Neutrino2004](mailto:Y.Suzuki@Neutrino2004)

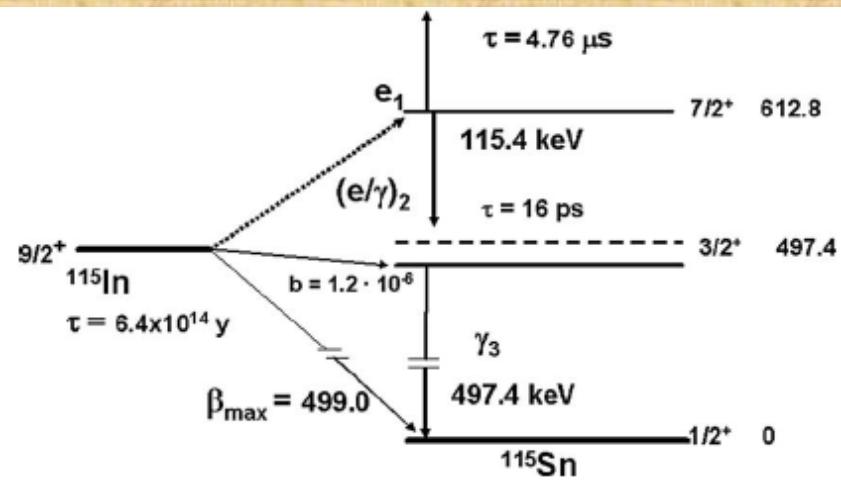


- mixing angle  $\theta_{12}$  is not well determined compared with  $\theta_{23}$  obtained by Atm.  $\nu$ .
- Survival probability could increase at 5MeV or less in case of LMA solution, and the value of probability depends on  $\theta_{12}$ .

pp/<sup>7</sup>Be solar neutrino spectrum gives us precise  $\theta_{12}$

# Capture of low energy solar neutrinos by $^{115}\text{In}$

R.S.Raghavan Phs.Rev.Lett37(1976)259



Nuclear Physics A 748 (2005) 333-347



## ● Advantage

- large cross section ( $\sim 640 \text{ SNU}$ )
- direct counting for solar neutrinos
- sensitive to low energy region ( $E_\nu \geq 125 \text{ keV}$ )
- energy measurement ( $E_e = E_\nu - 125 \text{ keV}$ )
- triple fold coincidence to extract neutrino signal from huge BG ( $e_1 + \gamma_2 + \gamma_3$ )

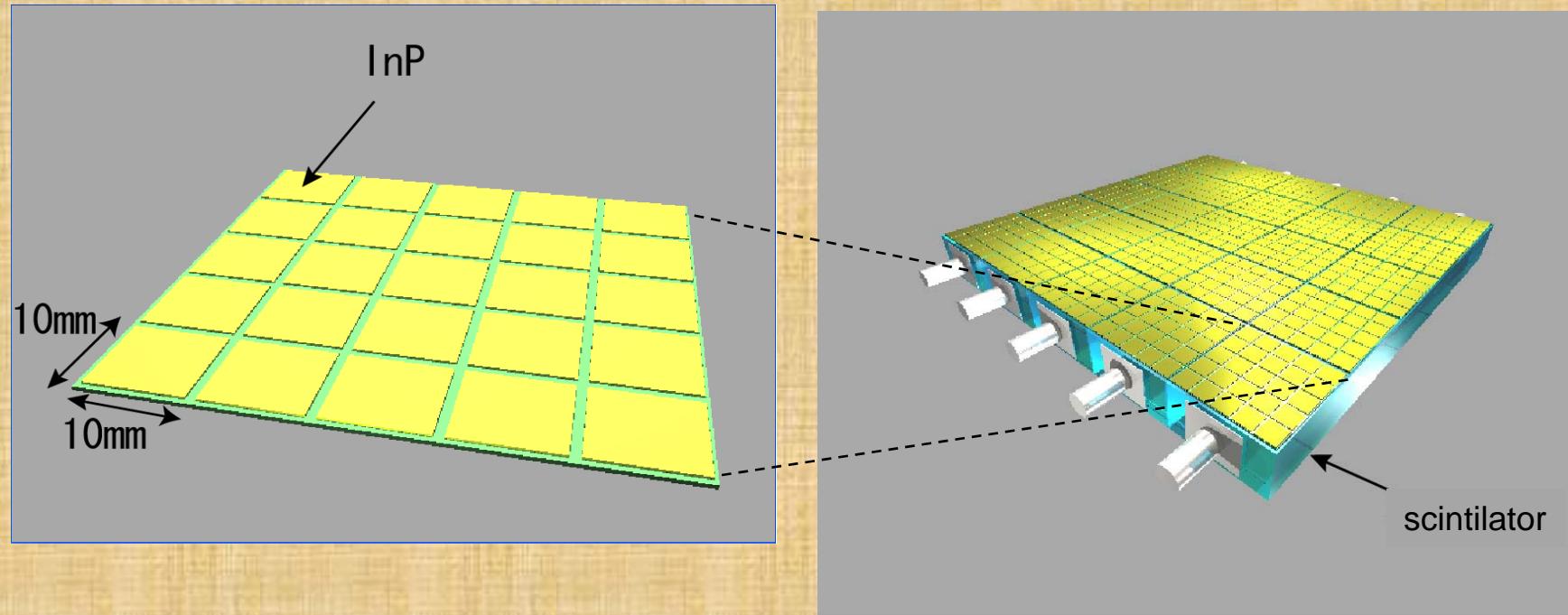
## ● Disadvantage

- natural  $\beta$ -decay of  $^{115}\text{In}$  ( $\tau_{1/2} = 4.4 \times 10^{14} \text{ yr}$ ,  $Ee \geq 498 \text{ keV}$ )
- possible BG due to correlated coincidence by **radiative Bremsstrahlung**

## Requirement for the detector

1. Good energy resolution : 10% (FWHM)
2. Fine segmentation ( $10^4$ – $10^5$ )
3. High efficiency  $\gamma$  detection

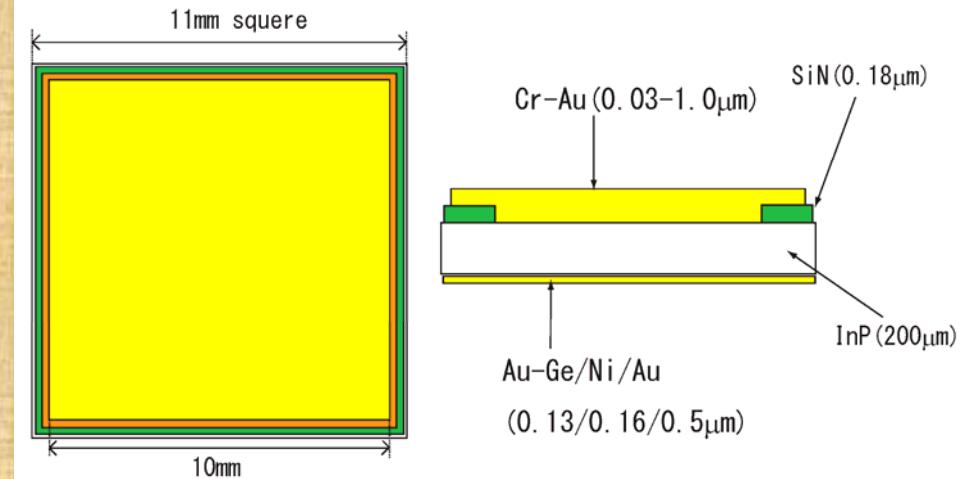
# Possible InP detector for solar neutrinos



- Multi-pixel structure for large area detector
- High Z scintillator surrounding InP detector detect  $\gamma$ s
- 4tons of  $^{115}\text{In}$  detector for low energy solar  $\nu$

Indium Project on Neutrino Observation for Solar interior  
(IPNOS) experiment

# Semi-insulating InP cell detector



## Mounted in vacuum chamber

- SI InP cell detector using VCZ-InP wafer (product of Sumitomo Electric K.K.)
- Cooled by dry-ice ( $T = -79$  degree)
- Response for gammas from radioactive sources

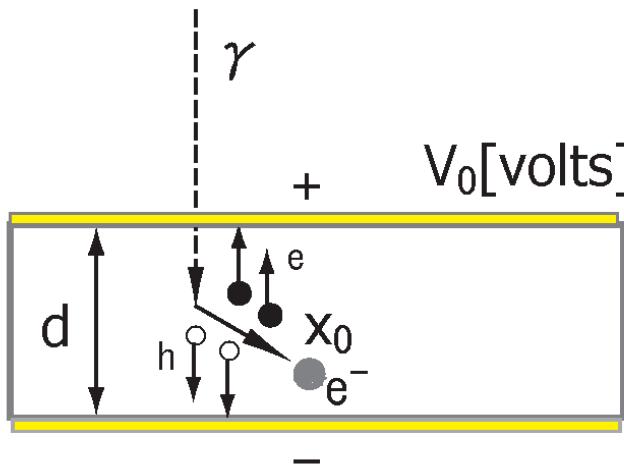
Surface size:

$10\text{mm} \times 10\text{mm} \times 0.2\text{mm}$   
 $(6\text{mm} \times 6\text{mm} \times 0.2/0.23/0.28/0.45\text{mm})$

Electrode :

- Ohmic contact
- evaporated Au base metal
- Insulator (SiN) to avoid leak current

# Principle of charge collection



$\mu$  : mobility [ $m^2V^{-1}s^{-1}$ ]

$v$  : carrier velocity [ $ms^{-1}$ ]

$E$  : electric field [ $Vm^{-1}$ ]

$d$  : thickness of SI InP

$x_0$  : range of electron

$$E = V_0/d \quad v = \mu E = \mu V_0/d$$

drift length :  $L_d = \tau v = \mu \tau V_0 / d$

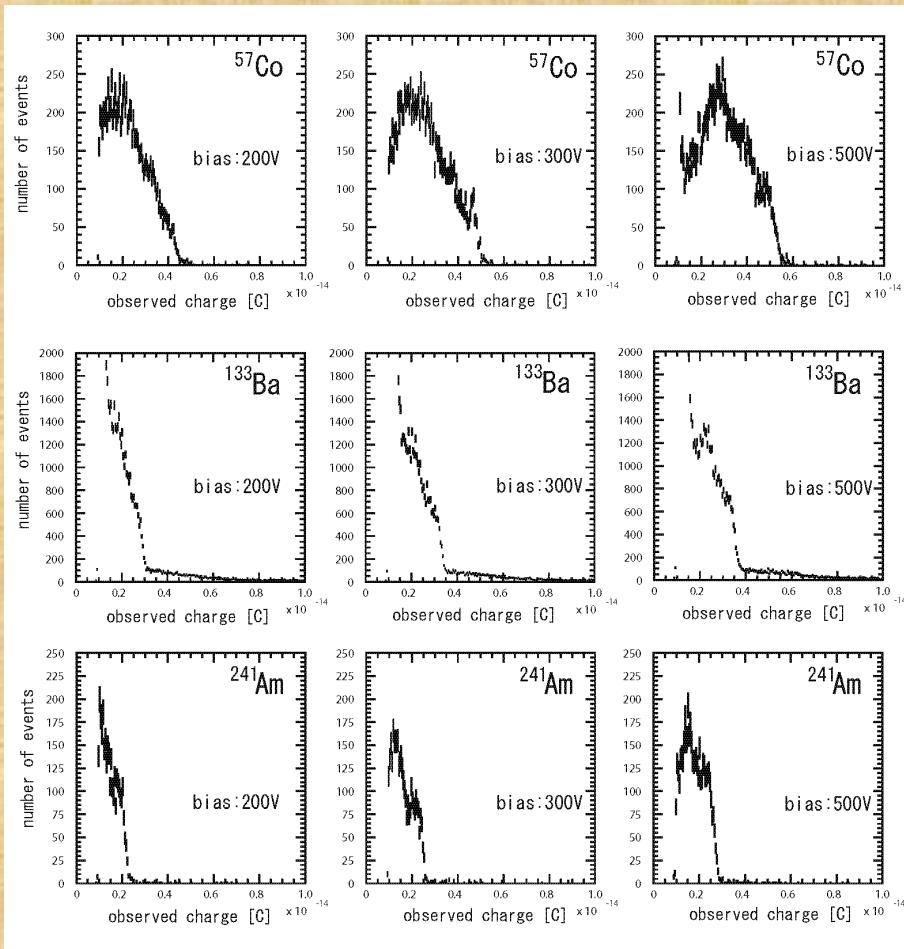
$\tau$  : carrier lifetime [s]

- **Induced charge** :  $dQ = (q/d) dx$
- Using Hecht formula,

$$Q = Q_0 \left\{ \left( \frac{L_e}{d} \right) \left( 1 - e^{-\frac{x}{L_e}} \right) + \left( \frac{L_h}{d} \right) \left( 1 - e^{-\frac{(d-x)}{L_h}} \right) \right\}$$

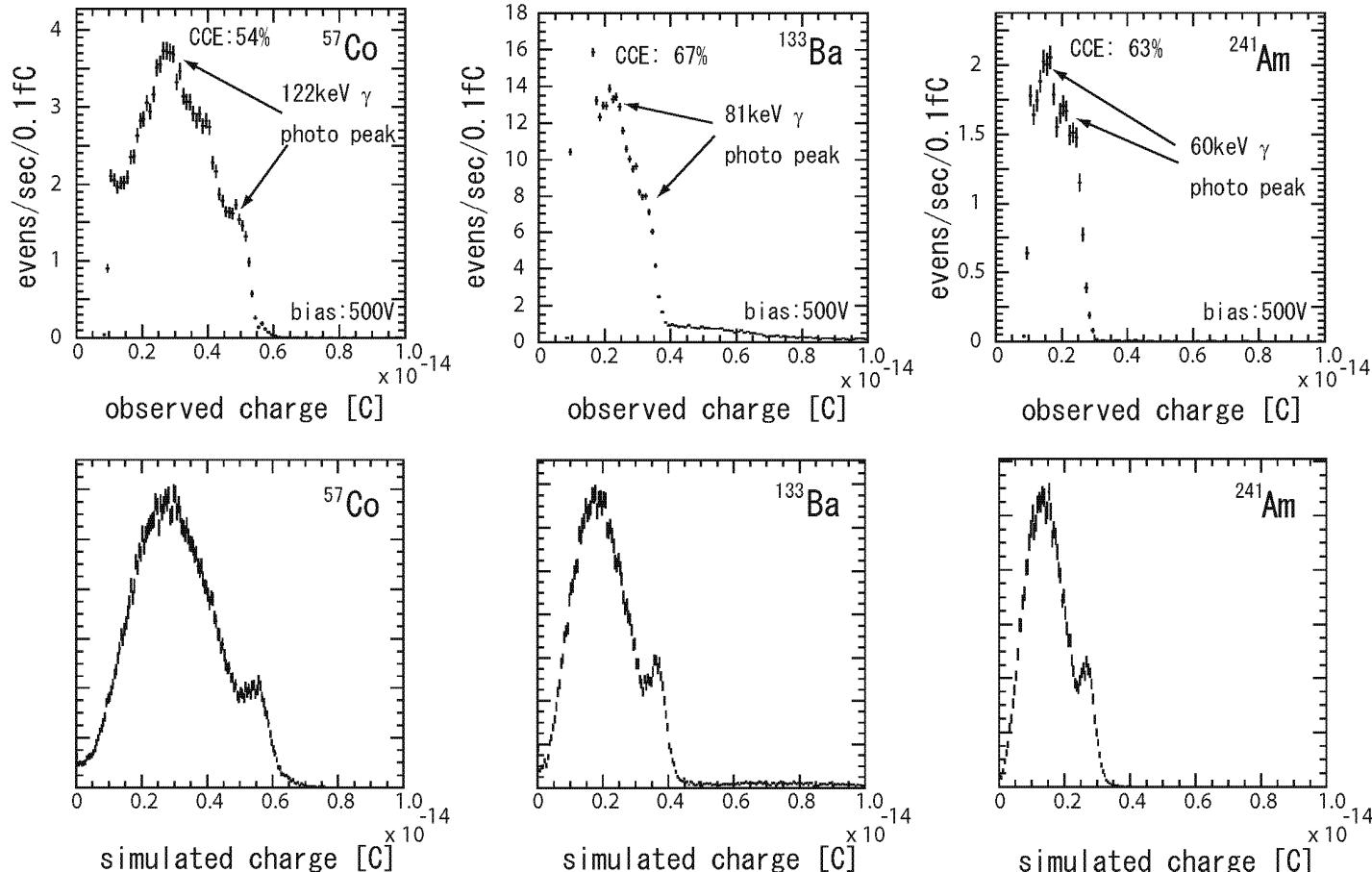
- For full collection : ( $L_e + L_h \sim d$ )  $Q = Q_0$

# $\gamma$ spectrum measured by InP detector



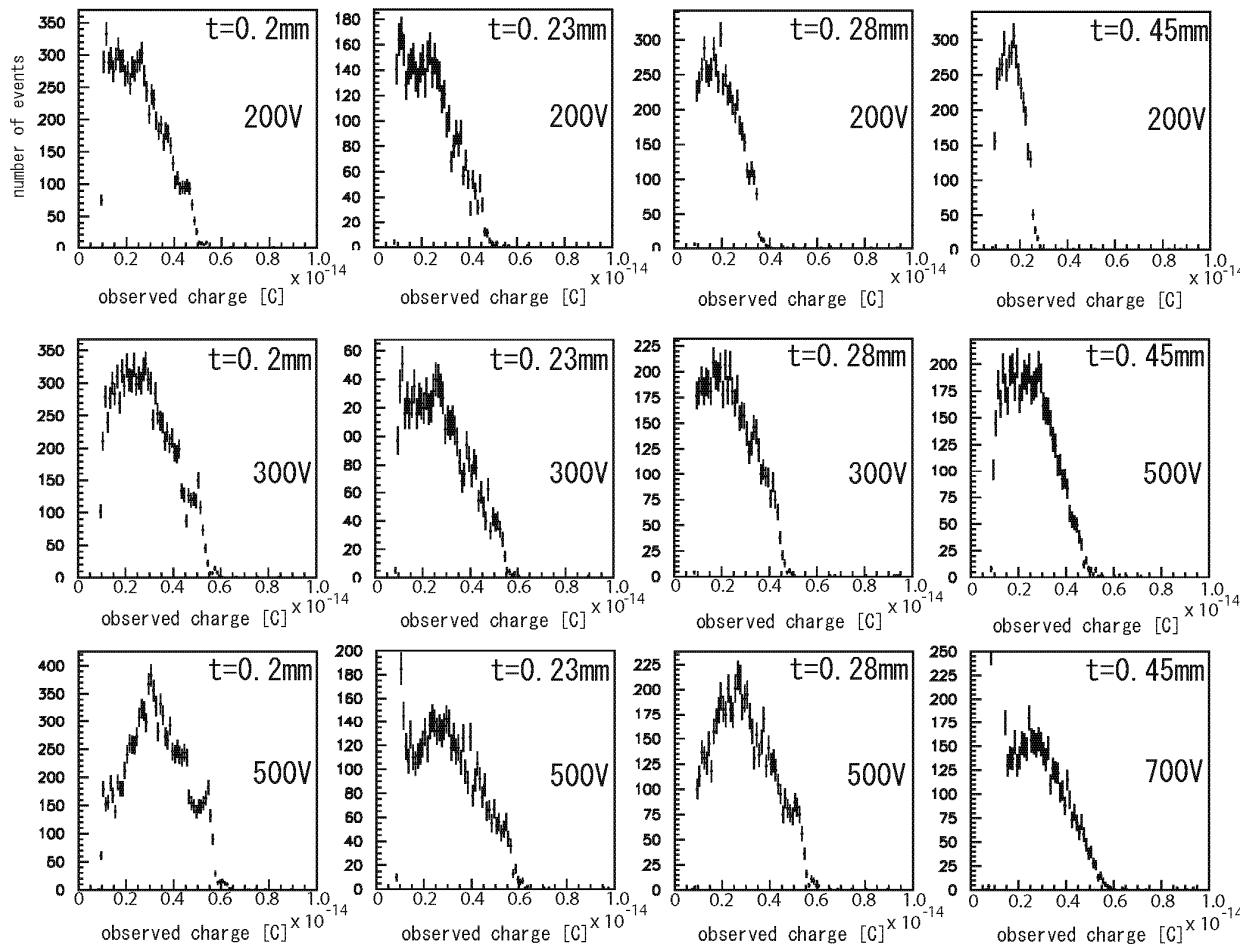
- InP detector should be cooled (-79 degree using Dry-Ice)
- Clear photo-peak was observed, but two peak structure
- Lower peak: induced charge generated by drift of carrier (electron and hole)
- Higher peak: full charge collection
- Energy of electron-hole pair production : **3.5eV**
- Energy resolution : **25%@122keV** for induced charge peak (intrinsic : 3%)

# Spectral shape and simulation



- Assuming,  $L_e \sim 200\mu\text{m}$  and  $L_h \sim 30\mu\text{m}$ , two peak structure could be reproduced by induced charge and full charge collection.

# Optimize detector thickness (using $^{57}\text{Co}$ )



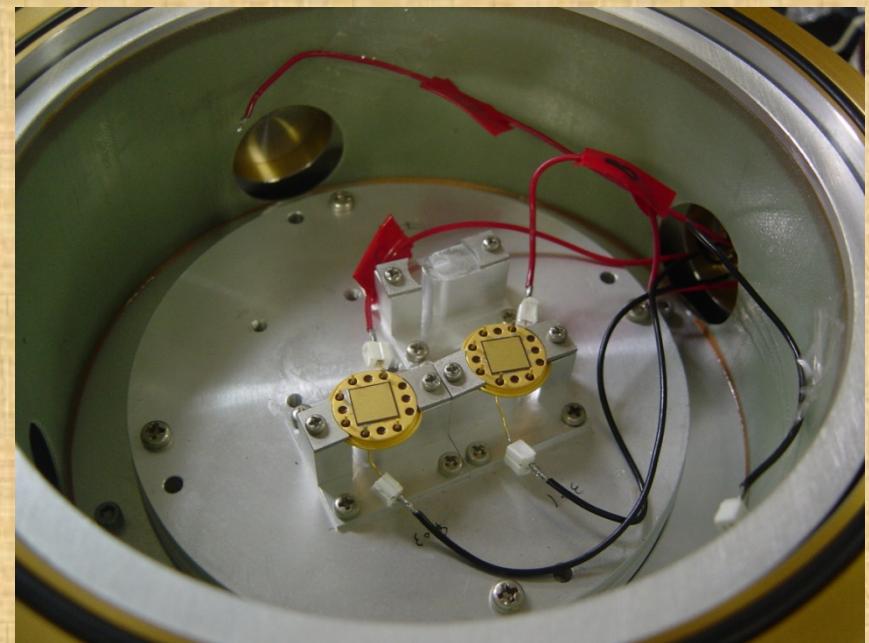
~ $280\mu\text{m}$  is best select for thickness

# Development for multi-pixel detector

**New dewar for parallel connection**

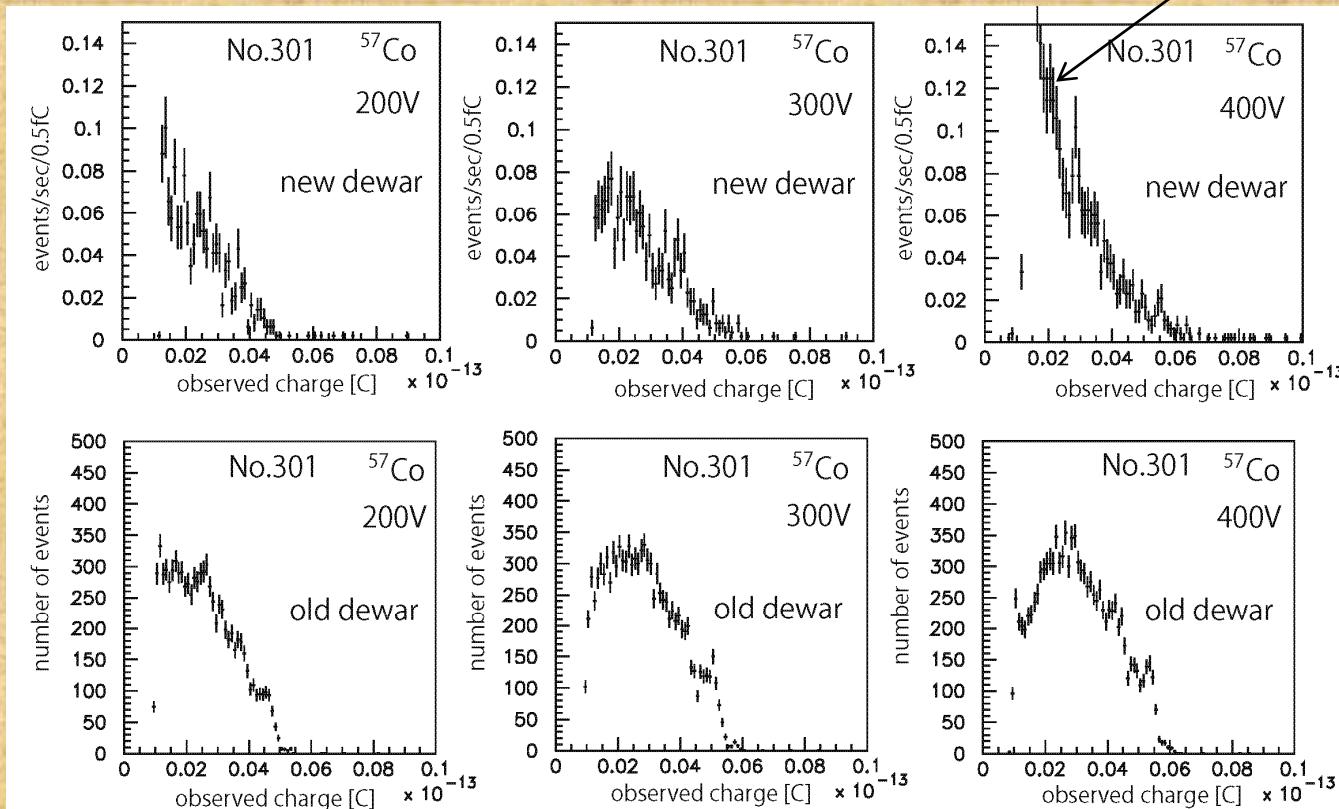


**Inside of new dewar for parallel connection**



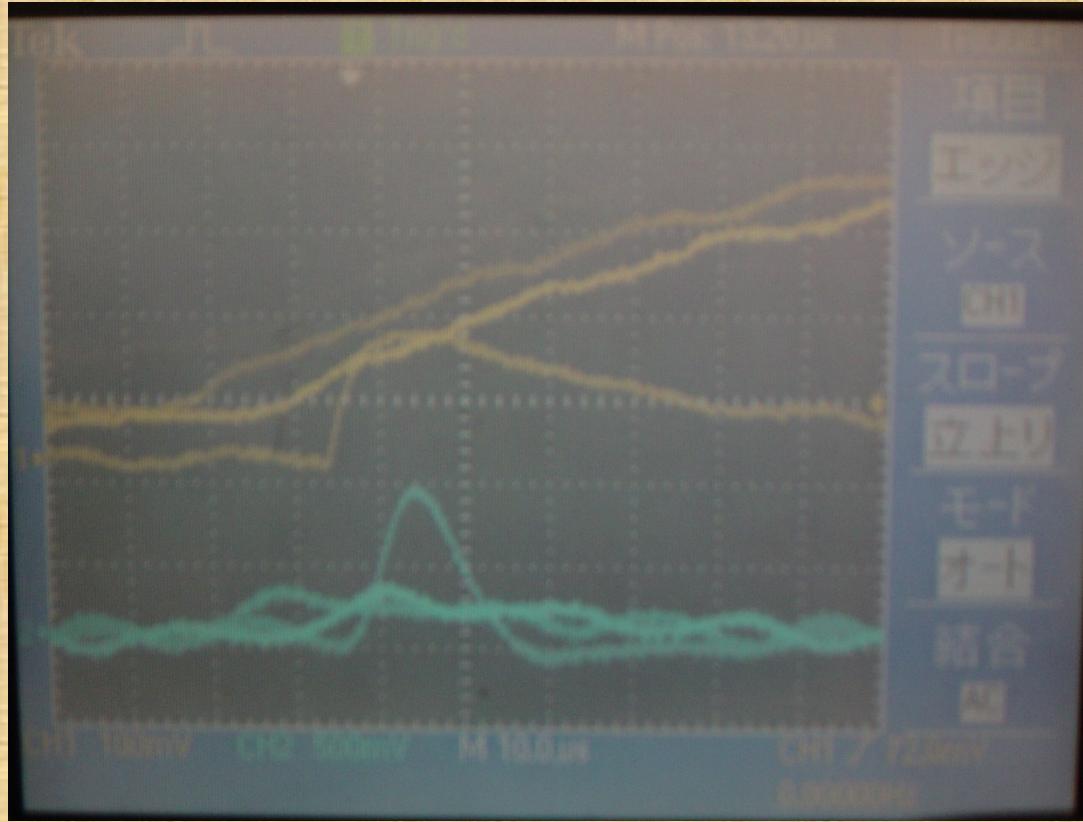
# Observation of $\gamma$ -rays for old and new dewar

## ■ $^{57}\text{Co}$ (Serial # 301)



Same peak and structure was observed.

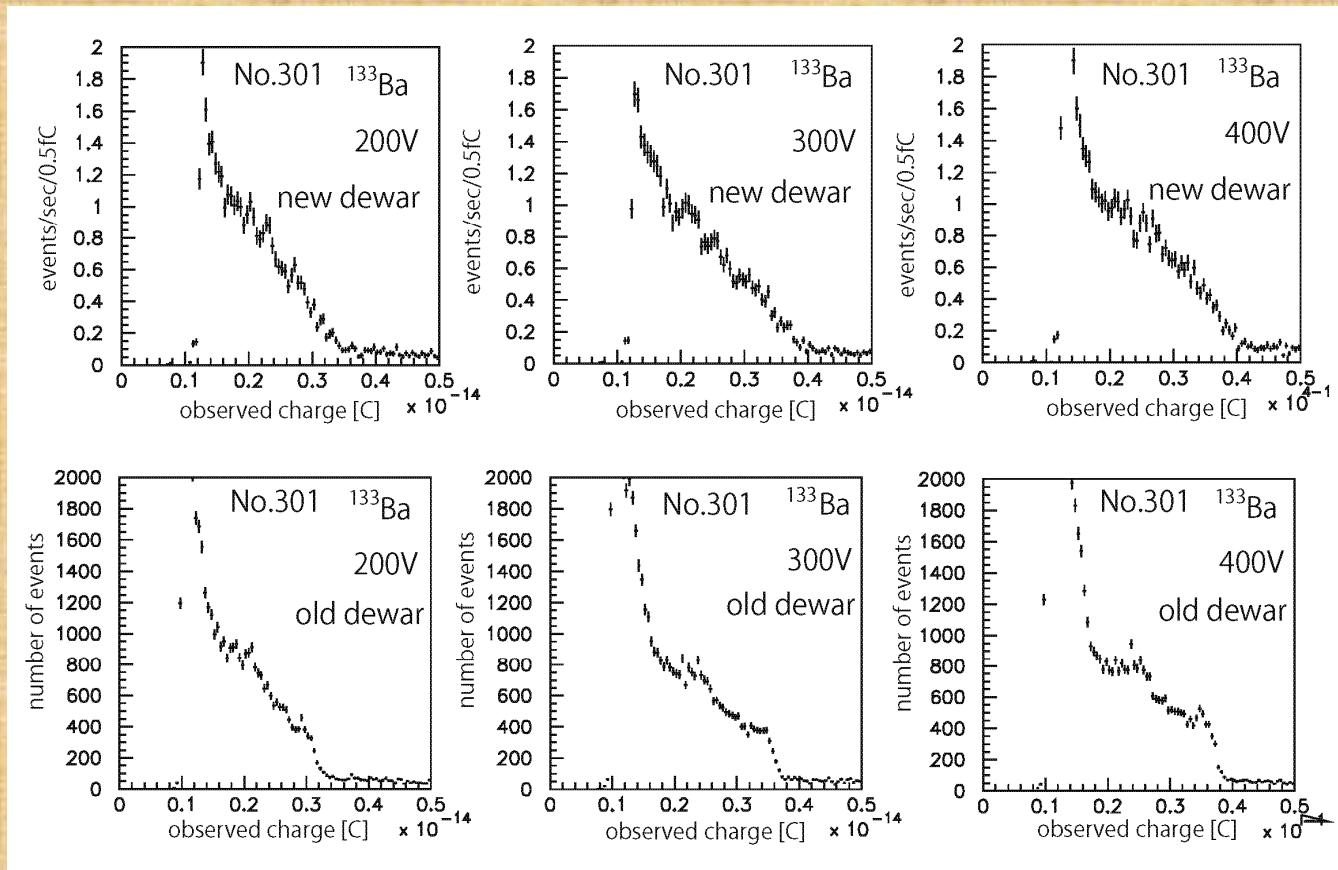
# Pulse-like noise



Unstable pulse-like noise sometimes happened in new dewar measurement.

# Observation of $\gamma$ -rays for old and new dewar

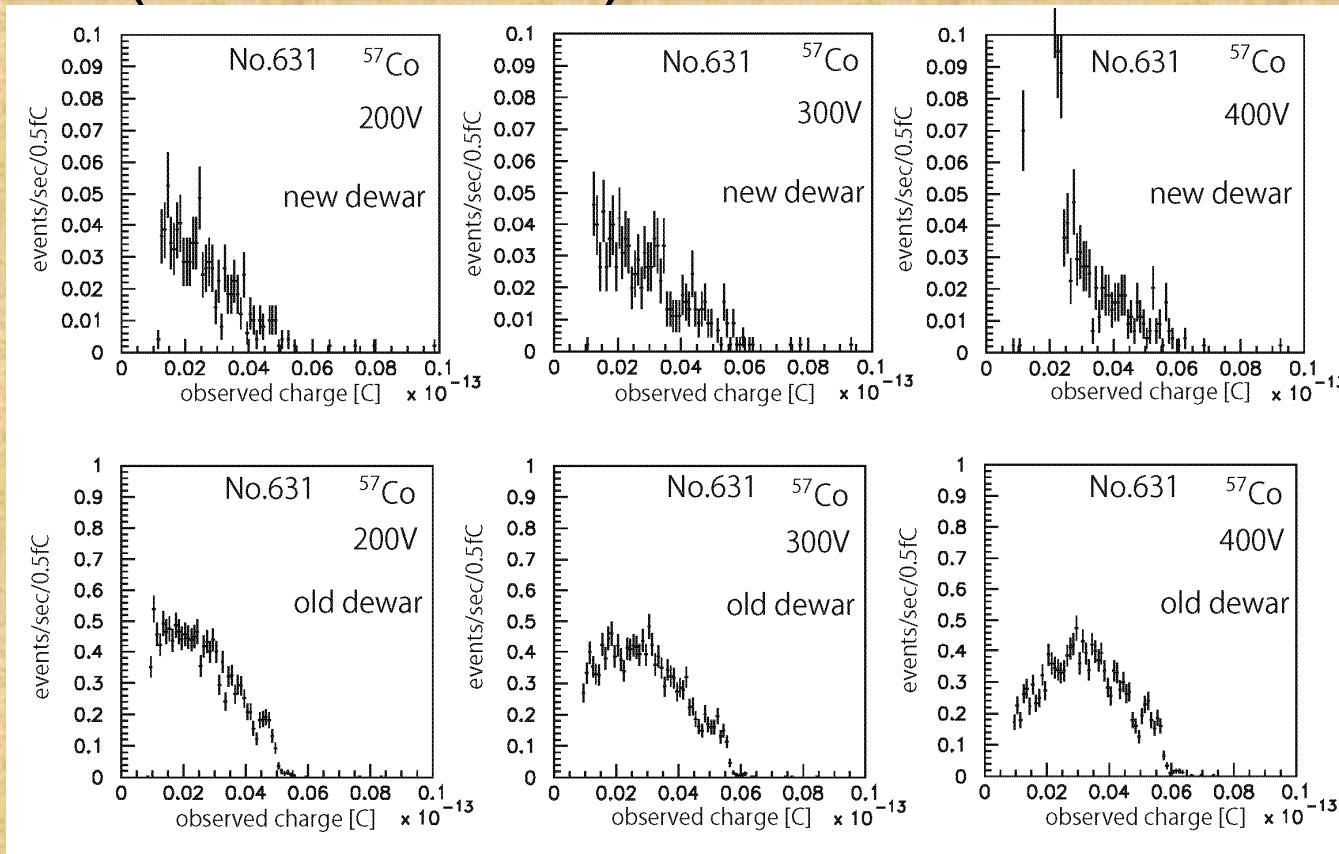
## ■ $^{133}\text{Ba}$ (Serial # 301)



Same structure, but worse resolution

# Observation of $\gamma$ -rays for old and new dewar

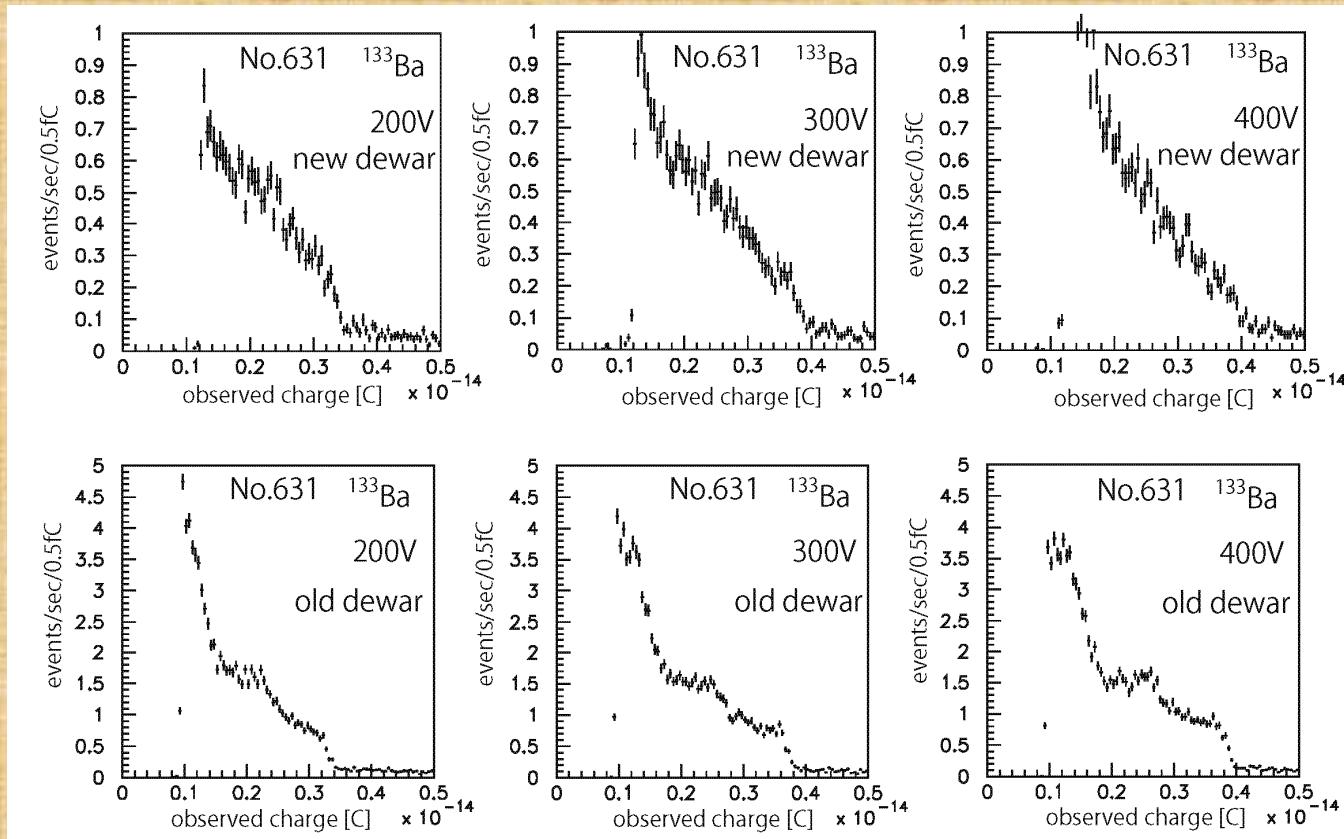
## ■ $^{57}\text{Co}$ (Serial # 631)



Also same structure, but more worth resolution

# Observation of $\gamma$ -rays for old and new dewar

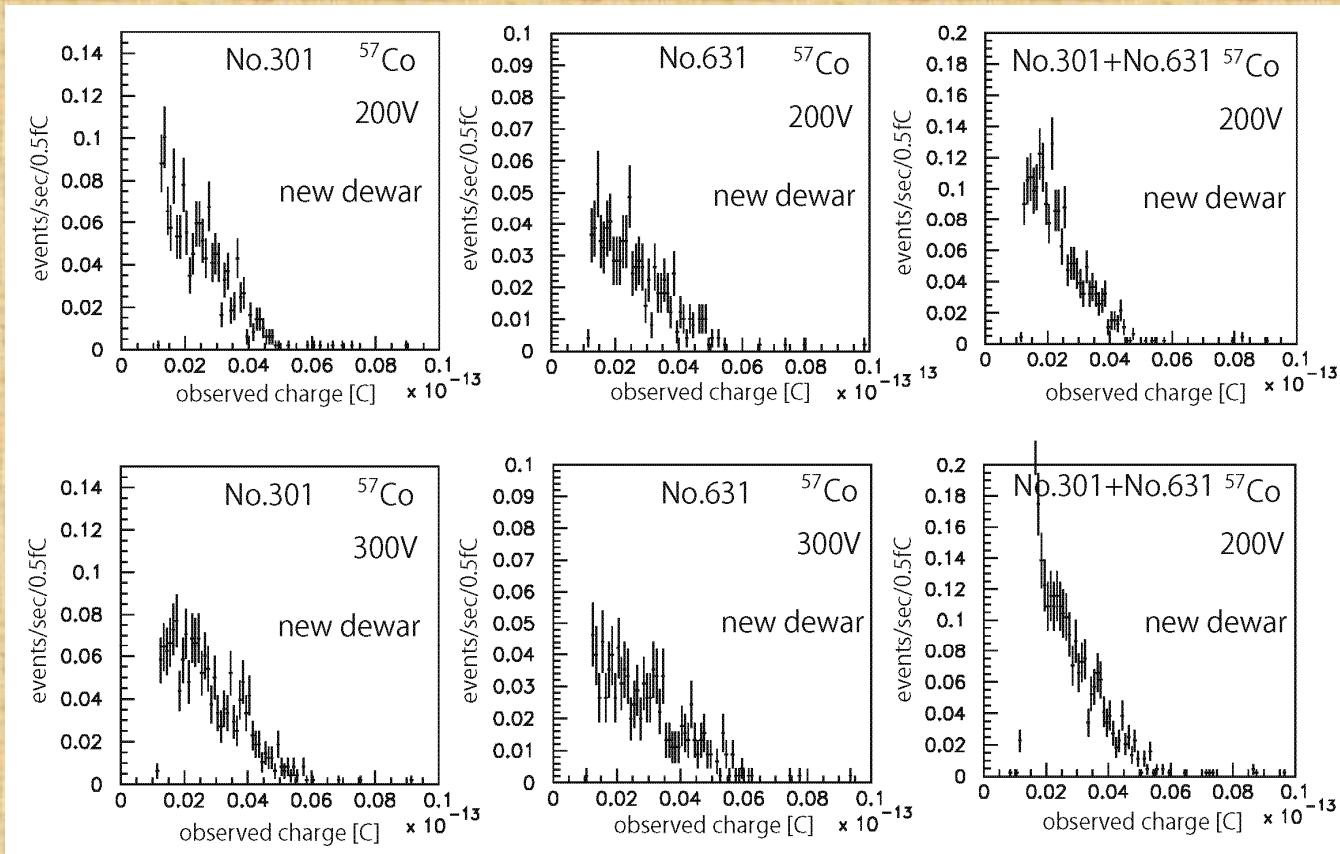
## ■ $^{133}\text{Ba}$ (Serial # 631)



Same structure, but more worse resolution

# Observation of $\gamma$ -rays for parallel connection

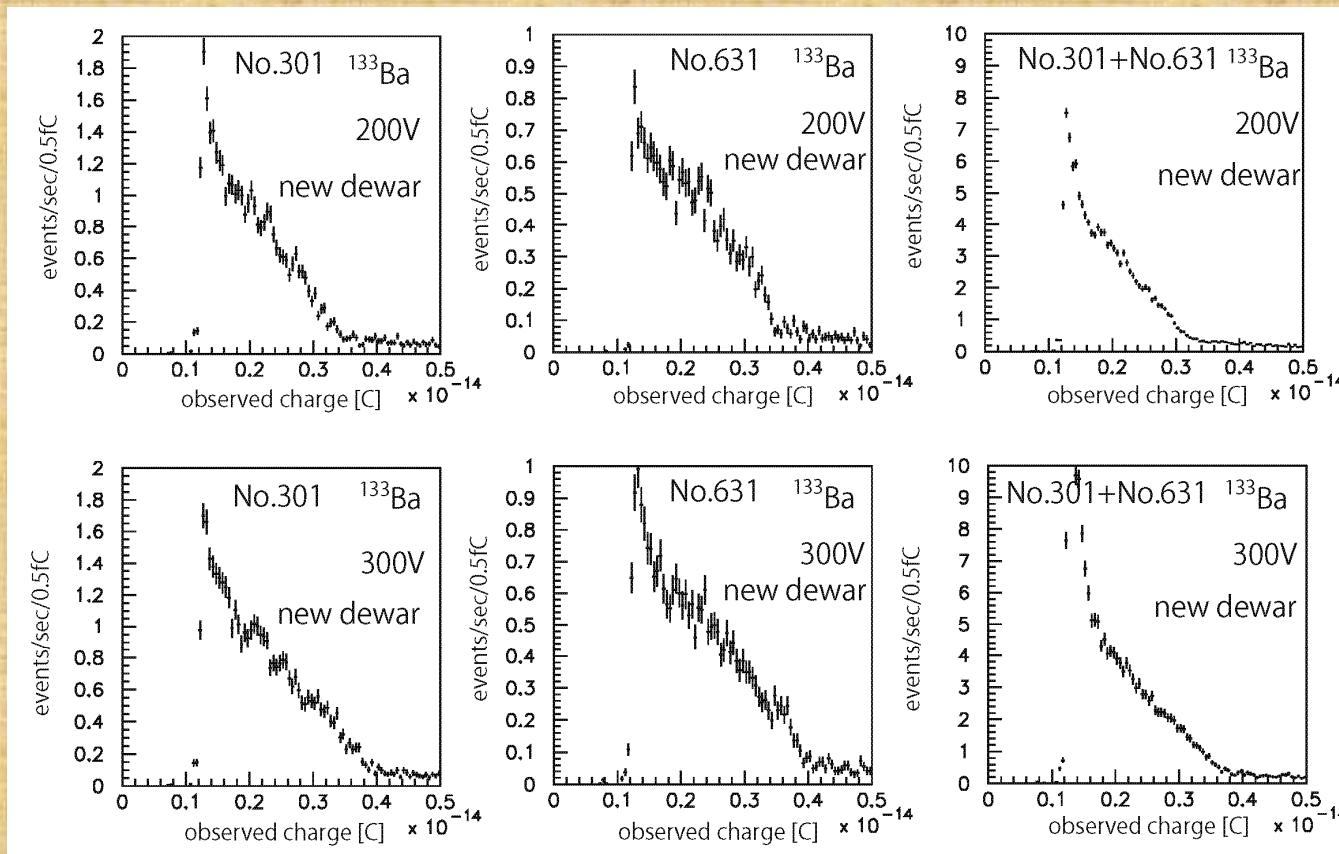
## ■ $^{57}\text{Co}$ (Serial #301 + Serial #631)



Events rate increased and same resolution

# Observation of $\gamma$ -rays for parallel connection

## ■ $^{133}\text{Ba}$ (Serial #301 + Serial #631)



Events rate increased and same resolution

# Conclusion

- Induced charge due to drift of carrier (electron and hole) generated by radiation.
- Charge distribution depends on thickness of detector.
- Parallel connection for multi-pixel detector was tested using NEW dewar.
- Charge distribution has same structure but worse resolution  **fix pulse-like noise problem**
- Events rate increased as expected (or more)
- More multi-pixel detector connection will be tested.

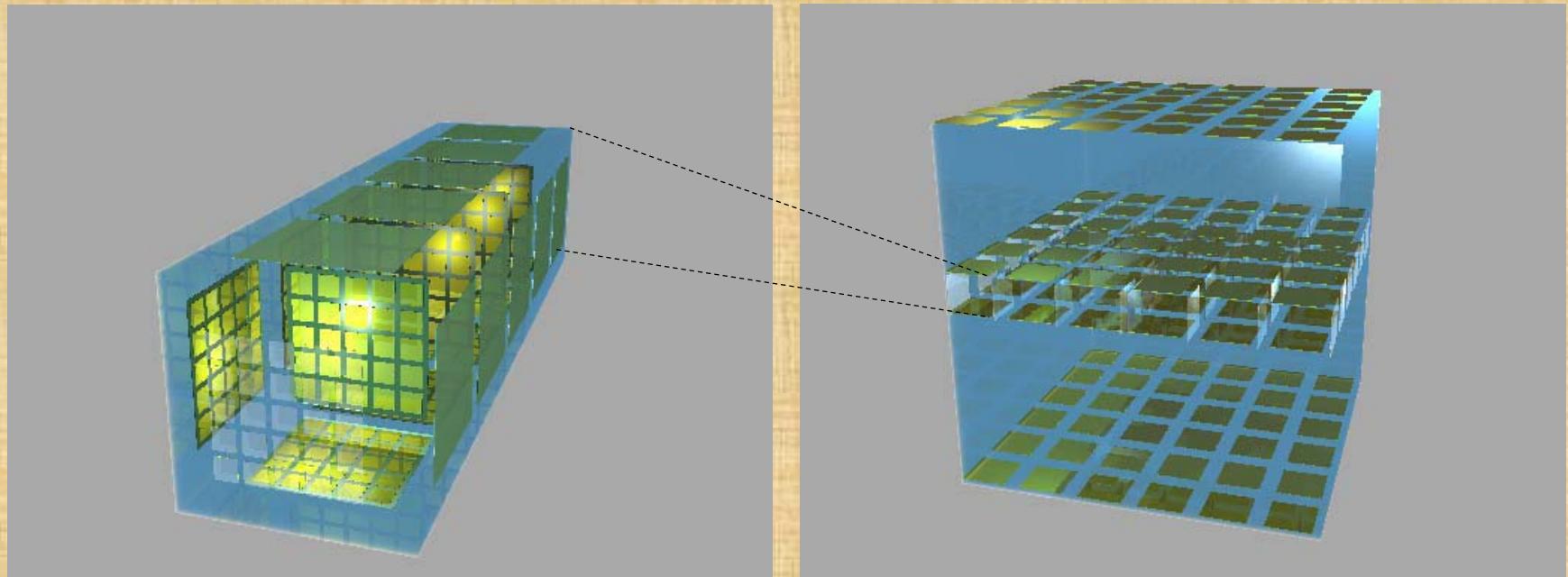
## IPNOS phase-I (10kg InP in LXe)

- ✓ Low background (&low temperature) environment inside of LXe
- ✓ A few events per year for pp solar neutrinos, but...
- ✓ Demonstrate actual performance for low energy solar  $\nu$

# New concept for IPNOS phase-I experiment

InP multi-pixel detector inside of Liquid Xenon.

30cm cubic chamber (like XMASS 100kg prototype) includes ~10kg InP detector



new detector to measure scintillation light from LXe

# Prototype of multi-pixel InP detector

